

INFN PHYSICS DEPARTEMENT OF FEDERICO II, NAPLES NOVEMBER 26, 2024





PARTGPATING SCHOOLS

LICEO TELESIO DI TELESE TERME

LICEO GALILEI DI NAPOLI



LICEO NOBEL DI TORRE DEL GRECO

La fisica dei raggi cosmici

La IISIca and Prof.SSA Laura Valore Università di Napoli Federico II INFN Sezione di Napoli 26 Novembre 2024

In cooperation with many networks and partners

IN

HATARE COSMIC RASS

Atomic nuclei produced in the cosmos that reach us on Earth

FONDOMEDETECTAND STUDY COSMIC RAYS?

DIRECT MEASUREMENT

THE FERMI GAMMA-RAY SPACE TELESCOPE

Direct measurement of cosmic rays involves capturing and analyzing high-energy particles using specialized detectors, such as cloud chambers or digital tracking systems. Indirect measurement relies on observing the secondary particles and radiation produced when cosmic rays interact with the Earth's atmosphere

INDIRECT MEASUREMENT

PIERRE AUGER OBSERVATORY ARGENTINA

Muons are subatomic particles similar to electrons but with a bigger mass, approximately 200 times that of an electron.

<text>

Muons penetrate any matter, better than any other known particles, except neutrinos.

They are classified as leptons and are unstable, decaying into other particles with a relatively short lifetime, about 2,2 microseconds. When subatomic particles like protons enter the Earth's atmosphere, they create a secondary shower of particles, including MUONS that make up a large portion of the background radiation present on the Earth's surface.

"MUON'S FLUX DEPENDS ON THE ZENITH ANGLE."

Muons that arrive perpendicular to the Earth's surface (arriving along the direction of the local zenith), travel the shortest distance in the atmosphere, while, for high angles of incidence, with respect to the zenith, the distance to travel is longer.

The greater the theta angle, the greater the distance traveled, and consequently, greater is the probability that the muons decay before reaching the Earth's surface, and the flux is lower.

GOSMIC RAYS CUBE

The Cosmic Ray Cube (CRC) is a portable device that detects muons. It measures particle flow and angles at different altitudes. It consists in four modules with two layers of scintillator bars arranged at right angles and has LEDs to show when particles pass through.

THE EXPERIMENT BEGINS

Each group of students counted the number of muons per minute, then reapeated the mesurements for different angles.

ENITH ANGLE	MEASUREMENT		
θ (°)	R1	R2	R3
0	50	48	49
15	48	48	47
27	39	38	50
45	39	39	34
59	19	19	19
87	10	10	8
θ (°)	R4	R5	R6
0	51		
15	48	49	52
27	50	39	39
45	34	32	34
59	19	22	22
87	8		

EXAMINE EXPERIMENTAL DATA

DATAANALYSIS

As expected, the number of muons per minute decreases, changing the zenith angle according to the $\cos^2 \theta$ law.

θ (°)	Teoratical	Sperimental	
0	1,00	0,99	
15	0,93	0,97	
27	0,79	0,85	
45	0,50	0,71	
59	0,27	0,4	
87	0,00	0,18	

SHARING THE RESULTS

VSIAND VS2 FROM LIGEO SCIENTIFICO TELESIO OF TELESE TERME

